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ION THRUSTER DIAGNOSTICS USING SPECTRAL LINE AMPLITUDES

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Abstract

The optical radiation from the plasma discharge of an electron bombardment mercury ion thruster was investigated. This work extends and refines earlier measurements, correcting certain ambiguities that arose in the analysis of line amplitudes. Using the measured ratio of the Hg I line amplitude at 3655 Å to that at 3650 Å, a theory incorporating a bimodal electron distribution (Maxwell electrons plus primary electrons) was used to obtain the average electron temperature and primary electron fraction in the thruster ion chamber. The electron temperature ranged from about 1.2 eV to 6.6 eV; whereas the primary electron fraction varied from zero percent to about 5 percent. These values depended upon the discharge voltage and the radial location of the measurement. The percentage of doubly ionized mercury produced in the chamber was also determined as a function of discharge voltage.

Introduction

In a previous paper (1) the authors presented the results of preliminary measurements of the optical spectra emanating from thruster plasmas. The present work extends and refines these measurements, correcting certain ambiguities arising in the earlier analysis. Basically, a method is described which permits evaluation of thruster discharge parameters from measured line amplitudes.

The observed optical spectral amplitude is a measure of the number of emitted photons of energy corresponding to the wavelength of the spectral line. The problem is to relate this amplitude (proportional to the number of transitions at a given wavelength) to the physical processes populating the energy level producing the photon emission (2). The following statements can be made about the physical nature of the discharge plasma within electron bombardment thrusters.

The discharge operates at low pressure (less than 10-3 torr) so that collisional broadening of spectral lines, collisional energy transfer between atoms, and three body volume recombination would be expected to have a negligible effect on the observed line amplitudes (3). With the possible exception of the resonant states, absorption should play a negligible role in excited state populations.

The ion chamber plasma, although essentially a steady-state discharge, was not in thermal equilibrium at the electron temperature. This can be shown by comparing calculated electron temperatures for an equilibrium plasma $^{(4)}$, $^{(4)}$, with Langmuir probe measured temperatures $^{(6)}$, $^{(7)}$, $^{(9)}$. The measured temperatures $^{(6)}$, with langmuir probe measured temperatures one to two orders of magnitude higher than those calculated from equilibrium theory. The discrepancy results from application of thermal equilibrium equations to nonequilibrium sustained plasmas $^{(10)}$.

Langmuir probe measurements of the energy distribution of such discharges (6) using a derivative method first described by Medicus (11) indicate that the electron distribution consists of nearly monoenergetic, or primary, electrons at an energy determined by the discharge chamber potential dif-

ference and a low temperature Maxwellian distribution. Although an attempt is made in Refs. 6 and 7 to account for this bimodal distribution, there exists no satisfactory theoretical explanation to date. In order to interpret the observed spectral line amplitudes in terms of the electron energy producing the emitted radiation, it is convenient to incorporate this bimodal electron distribution in a theoretical model of the discharge plasma. Because, in principle, both primary and Maxwellian electrons can contribute to excitation and ionization of mercury atoms, a theoretical model of the discharge plasma should incorporate excitation and ionization by both groups.

Theory

The rate of populating the $j^{ ext{th}}$ excited state due to collisional excitation from the ground state is given by

$$\dot{N}_{j} = N_{o}[N_{m}S_{j}(T_{e}) + N_{p}S_{j}(E_{p})]cm^{-3} sec^{-1}$$
 (1)

where $S_j(T_e)$ is the Maxwell averaged excitation coefficient for the j^{th} state and $S_j(E_p)$ is the excitation coefficient averaged over the primary electron distribution (represented as a delta-function at energy E_p).

function at energy E_p). The number of $j \rightarrow k$ transitions is then given by the steady-state relation

$$N_{jk} = N_{jk} \overline{A}_{jk} cm^{-3} sec^{-1}$$
 (2)

where \overline{A}_{jk} is the relative transition probability for the $j \to k$ transition, $\overline{A}_{jk} = A_{jk} / \sum_{k} A_{jk}$. Equa-

tion (2) can be rewritten with the aid of Equation (1).

$$N_{jk} = N_{oN_m}[1 + \phi X_{jk}(T_e, E_p)]S_{jk}(T_e)$$
 (3)

where

$$\phi = \frac{N_{p}}{N_{m}} \tag{4}$$

$$X_{jk}(T_e, E_p) = \frac{S_{jk}(E_p)}{S_{jk}(T_e)}$$
 (5)

and $S_{jk} = \langle Q_{jk} V_e \rangle$, the brackets denoting averaging over the appropriate electron distribution. The quantities Q_{jk} are related to the excitation cross sections Q_j by the relation

$$Q_{jk} = Q_j \overline{A}_{jk}$$
 (6)

They are the optical excitation functions for producing the observed spectral line corresponding to the $j \rightarrow k$ transition. The most recent measure-

ments of optical excitation functions for mercury are those of Anderson, et al. (12). A complete compilation of available mercury atom optical excitation functions can be found in Ref. 13. The number of transitions given by Equation (2) is proportional to the spectral line amplitude N_{λ} measured by the photodetector of the monochromater.

In order to obtain the electron temperature and primary electron fraction, it was necessary to ratio two line amplitudes. Examination of the observed Hg I mercury spectrum resulted in the selection of spectral lines at 3650 Å $(6^3D_3-6^3P_2)$ and 3655 Å $(6^3D_2-6^3P_2)$. Population of the 6^3D_3 level from higher lying states by cascading is negligible in the electron energy range of interest (12) and there is no singlet mixing of this state. The $63D_2$, however, mixes strongly with the $61D_2$ to produce an optical excitation function that is characteristic of the singlet state (12). The excitation cross sections for these two lines are thus sufficiently different in their dependence on electron energy that the line ratios vary significantly. Thus a method based on that described in Refs. 10 and 14 was used to obtain electron temperature and primary electron fraction.

From Equation (3) the ratio of the number of 3655 Å transitions to 3650 Å transitions is given by

$$\frac{N_{3655}}{N_{3650}} = \begin{bmatrix} 1 + \phi X_{3655}^{(T_e, E_p)} & S_{3655}^{(T_e)} \\ 1 + \phi X_{3650}^{(T_e, E_p)} & S_{3650}^{(T_e)} \end{bmatrix}$$
(7)

The variations of the appropriate excitation coefficients with electron temperature or primary electron energy are given in Figs. 1 and 2.

In order to use Equation (7) to obtain primary electron fractions ϕ and electron temperatures $T_e,$ it was necessary to choose a value for the primary electron energy $E_p.$ Previous studies of the discharge of such thrusters (15,16) indicated that the primary electron energy can be related to the discharge chamber potential difference by the expression

$$E_{p} = \Delta V_{I} - V_{P} \tag{8}$$

where Vp is the plasma potential in the cathode-magnetic pole piece region. In the present study this potential, Vp, was taken to be 16 volts, based on electric probe measurements reported in Ref. 16. As noted in this reference, however, Vp is dependent on cathode geometry, and thus the value of 16 volts is not universal.

A qualitative estimate of the sensitivity of the transition number ratio $\rm N_{3655}/\rm N_{3650}$ to primary electron energy (and thus the value of Vp used in Eq. (8)) is shown in Fig. 3. Here the left side of Equation (7) was plotted against Ep at two electron temperatures and three values of primary fraction. It was evident from such plots that the transition number ratio became less sensitive to changes in primary energy as the primary fraction decreased and the electron temperature increased. Thus, an error in the selection of Ep would have the largest effect at the lower electron temperatures and high primary fraction.

In summary, the theoretical model developed herein incorporates excitation by both Maxwellian and primary electrons. This represents a convenient model having simple features that allow for straightforward interpretation of the optical radiation from the thruster discharge chamber. Recent Langmuir probe measurements indicate that such a distribution does indeed exist in thruster discharges. The analysis uses measured spectral line amplitudes to calculate the electron temperature and primary electron fraction in the discharge chamber plasma of a bombardment thruster. Once these parameters are determined, other properties of the discharge such as ion fractions and density variations can be studied.

Experiment

The preceding theory was used to interpret the results of spectroscopic measurements obtained on a 30-cm diameter hollow cathode bombardment thruster. Detailed descriptions of such thrusters, shown schematically in Fig. 4, can be found in Refs. 17 and 18. The hollow cathode serves as the source of ionizing electrons which are contained in the discharge chamber by means of field-shaping permanent magnets. Propellant utilization, uncorrected for doubly charged mercury ion production, ranged from about 80 percent to about 100 percent. The extraction grid open area was about 72 percent.

An Ebert mounted, plane grating monochromator with a 0.5-meter focal length was used to study discrete spectral line amplitudes. The instrument dispersion was 16 Å/mm. The plane grating could be rotated to cover the wavelength range from 2000 Å to 8000 Å. The recorded photomultiplier signal was corrected for the spectral response of the optical system. This response was determined using a radiance standard calibrated by the National Bureau of Standards.

The thruster was operated in the same 7.6-meter diameter by 18.3-meter facility described in Ref. 1. There was a question as to how much of an error would be introduced as a result of viewing the ionization chamber through the exhaust beam. This downstream radiation was measured by locating the monochromator normal to the thruster axis and viewing a region just downstream of the thruster. Such measurements indicated that the downstream radiation was about two to three orders of magnitude less than that produced in the discharge chamber and could thus be neglected.

The monochromator was used to obtain radial profiles of the 30-cm diameter thruster. The procedure consisted of setting the monochromator on the peak of a given spectral line. A motor driven 25 micron horizontal slit was then used to scan the image on the vertical entrance slit of the monochromator. In this manner the area of thruster cross section observed represented about 0.005 percent of the total thruster cross sectional area. For measurements of the thruster discharge chamber plasma, profiles were obtained along the radial direction away from the neutralizer location as shown in Fig. 5.

Results

Spectral line amplitudes were obtained for discharge chamber potential differences ΔV^{\intercal} of 33, 38, 43, and 58 volts. The discharge current was maintained at 8 amperes, and the thruster beam current was constant at 1.5 amperes. At each discharge voltage the ratio of the number of 3655 Å to 3650 Å transitions was calculated at four radial positions. These ratios are shown graphically in Fig. 6. Here the measured ratio N_{3655}/N_{3650} is

plotted at the four radial positions. Of particular interest is the minimum occurring at a primary electron energy of 22 eV (ΔV_{I} = 38 V) at all radial positions except that furthest from the thruster axis. It will be shown that such a minimum can be explained by the existence of nonrandomized primary electrons in the plasma discharge.

In Fig. 7, Equation (7) is plotted against electron temperature for the four primary electron energies studied. Several curves corresponding to different values of primary electron fraction ϕ ranging from 0 to 0.1 are presented. The problem is to correlate these curves with the experimental curves of Fig. 6. In this way, the N_{3655}/N_{3650} ratio, measured as a function of primary electron energy, can yield the electron temperature and primary electron fraction corresponding to a particular primary energy. Note that in Fig. 7, curves of $\phi > 0$ also show minima, analogous to the minima of the data curves of Fig. 6 at a primary electron energy of 22 eV. It is this fact which suggests the procedure to be used to interpret the spectral amplitude data. In reducing the data, use is made of the following conditions:

Condition (1): Increasing the primary electron energy increased the available discharge power per atom, presumably increasing the electron temperature. Thus the electron temperature obtained from Fig. 7(d) would be greater than the temperature obtained from Fig. 7(a).

Condition (2): Because an increase in primary electron energy increases the mean relaxation time for primary electrons (19), it was assumed that the primary electron fraction did not decrease with increasing primary electron energy. The average primary fraction obtained from Fig. 7(d) would thus be greater than the average fraction obtained from Fig. 7(a).

Condition (3): For values of E_D up to 22 eV and for normalized radii less than 0.75, the measured N₃₆₅₅/N₃₆₅₀ ratio (Fig. 6) decreased with increasing electron energy. Also, the theoretical N₃₆₅₅/N₃₆₅₀ ratios of Fig. 7 decreased with increasing electron temperature for temperatures up to about $^{\text{h}}$ eV. Thus for E_D \leq 22 eV, portions of the ϕ -curves of Figs. 7(a) and (b) to the right of the minima locus (unused region) can be neglected. Similarly, for E_D > 22 eV, portions of the ϕ -curves of Figs. 7(c) and (d) to the left of the minima locus (unused region) are neglected.

The allowable range of values for primary electron fraction at different radial locations was obtained using Condition (2) and the following procedure. For a given radial location (e.g., r = 0) and primary electron energy equal to 22 eV, a line was drawn parallel to the abscissa of Fig. 7(b) so as to intersect the ordinate at a value equal to the measured N_{3655}/N_{3650} ratio (at r = 0, the ratio was 0.065, from Fig. 6). This line is tangent to a particular o-curve at its minimum, and thus intersects the locus of minima at a particular value of ϕ and T_e . (For r = 0, the measured N_{3655}/N_{3650} line was tangent to the $\phi = 0.03$ curve at a temperature Te of about 3.5 eV.) This value of $\,\varphi\,$ represents an upper bound of the $\,\varphi\,$ (and also T_e) range at the particular radial location and for primary electron energies from 17 eV to 22 eV. Lower bounds on ϕ and $T_{\mbox{\scriptsize e}}$ for this primary electron energy range and for a particular radial location were obtained from the measured N_{3655}/N_{3650} ratio at 17 eV and represented by a horizontal line overlayed on Fig. 7(a). The intersection of this line with the $\phi\text{-curves}$ to the left of the minima and the assumption that for tenuous plasmas, the electron temperature is generally expected to exceed about 1 eV(10) served to obtain these lower bounds (Thus, for r = 0, where N3655/N3650 = 0.074 at $E_p=17$ eV, the lower bound on ϕ from Fig. 7(a) was found to be about 0.01 at an electron temperature of about 1.2 eV.) The lowest lower bound on ϕ was obtained at a normalized radius of 0.75 and was $\phi=0$.

For Ep > 22 eV, the procedure used to obtain the appropriate ranges of electron temperature and primary electron fraction was as follows: The maximum primary electron fraction was obtained from the intersection of the horizontal line corresponding to the measured N₃₆₅₅/N₃₆₅₀ ratio with the curves of Fig. 7(d). The largest value of this ratio (equal to 0.089) was measured at $\rm r=0$, for which the largest value of ϕ intersected was about 0.05. Thus, for the lower primary energy of 28 eV (Fig. 7(c)), the intersection of measured N₃₆₅₅/N₃₆₅₀ values with curves of ϕ > 0.05 were disregarded in accordance with Condition (2).

The minimum value of ϕ at a given normalized radius and for $E_{\rm p} > 22$ eV could not be less than the value of ϕ obtained at 22 eV, again in accordance with Condition (2). Thus for r=0 and $E_{\rm p} > 22$ eV, the range of ϕ was defined to be 0.03 $\leq \phi \lesssim 0.05$. For $E_{\rm p} \leq 22$ eV and r=0, the range was 0.01 $\leq p \leq 0.03$. Using these ϕ ranges and the appropriate curves of Fig. 7, the range of electron temperature for r=0 and different values of $E_{\rm p}$ were also determined. Analogous procedures to those just described were used to obtain the ranges of ϕ and $T_{\rm e}$ at the other three radial locations studied.

The results of the analysis are summarized in Figs. 8(a) through (d). Here the variation in the range of electron temperature with primary electron energy is presented for four radial positions. The boundaries of the darkened regions were determined by the range of primary electron fraction. The values of electron temperature obtained were in general agreement with Langmuir probe measurements of hollow cathode thruster discharges $^{(9)}$. The fact that the largest values of ϕ and $T_{\rm e}$ were obtained on the thruster axis, whereas the smallest values of these parameters were obtained at three-quarters radius, is consistent with recent measurements of a primary electron region centered about the thruster axis $^{(20)}$.

Effects of metastables. - For the bombardment thruster discharge chamber plasma, the contribution to excited states by collisional excitation from metastable states can be shown to be much less than the direct excitation from the ground state and could thus be neglected. The two important metastable states in mercury are the $6^{3}P_{0}$ and $6^{3}P_{2}$ states. The ratio between the rate of collisional excitation of the 63D3 state of mercury from the 63P_O metastable level to the rate of collisionally populating the 63D3 from the ground state was calculated to be about 0.005 over an electron temperature range of 4 to 7 eV. The corresponding ratio for excitation from the 63P2 metastable state was about 0.027 over the same electron temperature range. These ratios were lower at electron temperatures below 4 eV. The calculations were based on the Gryzinski theory(21,22) and the assumption of a pure Maxwellian electron distribution. Experimentally, the primary electron density was generally less than 5 percent of the Maxwellian

electron density. This fact, in conjunction with the fact that excitation coefficients for primary and Maxwellian electron excitation are of the same order of magnitude (Figs. 1 and 2) implies that these estimates of the metastable level contribution to Equation (1) are reasonable. Thus, the exclusion of the metastable contribution to the collisional excitation of the 63D states was justified.

Ion excitation. - For ion excitation the spectral amplitude at 3984 Å, corresponding to the $6^2P_3/2$ - $6s^2$ $^2D_5/2$ transition in Hg II, was used as the diagnostic line. This line produced the strongest signal in the near visible Hg II excitation spectrum. The calculated excitation cross section for the resonant $6^2P_{3/2}$ level is given in Fig. 9. From the work of Penkin, et al.(23) one would expect cascading to contribute to the population of this level. Available data, however, are insufficient to estimate the extent of the cascade contribution. Also shown is the calculated level excitation cross sections for the $5d_{5/2}^{8}$, $3/26s^{2}$ (J = 4) Hg III line at 4797 Å. This line was used to obtain the fraction of Hg III in the discharge plasma. These cross sections are plotted as functions of electron energy. Excitation from the ground state of the ion (single ion for the 3984 A line and double ion for the 4797 A line) was assumed. From comparisons between such calculated cross sections for helium and experimentally measured helium optical excitation functions (22) the calculated cross sections for mercury are expected to be within an order of magnitude of actual cross sections.

<u>Density variations</u>. - The variation of the ion to neutral fraction was studied as a function of primary electron energy at different radial locations in a 30-cm hollow cathode thruster. Using the method to obtain Equation (7) measured amplitudes at 3984 A for Hg II and 3650 A for Hg I were ratioed to obtain the expression

$$\frac{N_{+}}{N_{0}} = \frac{N_{3984}}{N_{3650}} \frac{1 + \phi X_{3650}(E_{p}, T_{e})}{1 + \phi X_{3984}(E_{p}, T_{e})} \frac{S_{3650}(T_{e})}{S_{3984}(T_{e})}$$
(9)

where
$$S_{3984}(T_e) = \left\langle Q_{6^2P_3/2} \right. v_e \right\rangle_m \overline{A}_{3984}$$
.

The quantity A3984 is the relative transition probability. Its value is unknown, but is expected to be substantially less than one. This is because the $6^2P_{3/2}$ upper state of the 3984 Å transition is a resonant state and thus would radiate to the ground state. The excitation coefficient ratios $\chi_{\lambda}(\mathbb{E}_{p}, \mathbb{T}_{e})$ have been defined earlier. Because π_{3981} is constant, the right side of Equation (9) yields the relative variation of N_{+}/N_{0} . The range of values of ϕ and T_{e} used were obtained from Fig. 8. The relative variation of $N+/N_0$ is given in Table I (third column) as a function of E_{p} and radius. The calculations indicate that the ion fraction N_+/N_0 did not vary appreciably with discharge voltage, showing a maximum at a primary electron energy of 22 eV for radial locations near the thruster axis.

The relative variation of $\ensuremath{\,\mathrm{N}_{+}}$ can be obtained from the relation

$$N_{+} = \left\{ \frac{N_{398h}}{\left[1 + \phi X_{398h}(E_{p}, T_{e})\right] S_{398h}(T_{e})} \right\}^{1/2}$$
 (10)

tion (10) and Equation (9), the relative variation of N_0 can also be obtained. These variations are shown in Fig. 10, in which an average of the normalized ion and neutral atom densities are plotted as functions of primary electron energy. Maximum densities occurred at the lowest energy (17 eV). For each radial location the data were normalized to unity at this energy. The marked decrease in both ion and atom densities can probably be attributed to the decrease in propellant flow required to maintain the constant beam current (1.5 amperes) with increased discharge chamber potential difference. This fact suggests the possibility of relating propellant flow changes to changes in atom and ion spectral amplitudes. Also noted was a leveling off of the neutral density at high discharge chamber voltages. This result is in agreement with a conclusion of Ref. 20, stating that the neutral density remains constant at high discharge power levels (high eV/ion).

Fraction of doubly charged ions. - Excitation of doubly-charged mercury ions at 1797 A were also measured in the 30-cm diameter thruster. An expression analogous to Equation (9) (with 4797 ${\rm \AA}$ substituted for 3984 ${\rm \AA}$) can be used to obtain N_{++}/N_0 . In this case, $\overline{A}_{4797} = 1$, because the upper state of this transition has only a single path for radiative decay $^{(24)}$. Thus the spectral amplitude ratio gives the fraction of N++ in the discharge directly. These fractions are given in the last column of Table I. The accuracy of these measurements depends markedly on the accuracy of the excitation coefficients used. Because no measured excitation functions for Hg III are available to the authors' knowledge, it is not possible at present to assess the accuracy of these calculations. It should be emphasized that the double ion fractions presented here cannot be directly compared with mass spectrometer measurements of double ions in the beam (25). This is because the fraction N_{+}/N_{0} is not known, and beam measurements yield N_{++}/N_{+} ratios.

It should be noted that excitation of Hg III at an electron energy of 17 eV was observed at all radial positions, although quite weakly. The existence of Hg III at such low electron energies is to be expected because the threshold for ionizing Hg II to Hg III from the $6^2\mathrm{D}_{5/2}$ metastable level is about 15 eV. In addition, the "tail" of the low temperature Maxwellian distribution of electrons can produce both ionization of Hg II and excitation of Hg III.

Concluding Remarks

Spectral amplitude measurements indicated that the discharge was a nonequilibrium sustained plasma. The energy distribution of electrons in the discharge of hollow cathode thrusters can be represented by a distribution of monoenergetic electrons superimposed on a Maxwellian electron distribution, as has been suggested by other workers. Spectral line amplitudes at 3655 Å and 3650 Å were ratioed to obtain electron temperatures and ratios of primary to Maxwellian electron densities at four primary electron energies and different radial locations. It was found that the electron temperature ranged from about 1.2 to 6.6 eV; whereas the primary to Maxwellian electron density ratio ranged from zero to about 0.05. These values depended upon primary electron energy and radial location in the thruster.

Normalized ion and neutral densities were found to decrease with increasing discharge voltage at constant emission current and beam current. It was presumed that this behavior was attributable to a decreasing propellant flow with increasing voltage. This decrease in flow was required in order to maintain constant beam current. The neutral density leveled off at high discharge voltages (high eV/ion) in agreement with earlier reported work.

The fraction of doubly charged mercury ions did not exceed about 2 percent of the atom density in the discharge plasma at a discharge voltage of 58 V. At lower discharge voltages this fraction did not exceed about 0.8 percent.

One of the primary purposes of this investigation was to determine the extent to which the optical radiation emanating from the ion chamber of electron bombardment thrusters could be used to study the discharges of such thrusters. The high sensitivity of the radiation output to changes in thruster parameters suggests potential engineering applications. For example, studies are presently underway to investigate the light output from the neutralizer region. Such studies should lead to a better understanding of the operation of this thruster component. It is possible that the light output could be used in a neutralizer controller. It is concluded that this study demonstrated the possibility for using spectroscopic techniques as diagnostic tools for thruster research.

Sym	bo	1	s

^A jk	transition probability for $\mathfrak{J} \rightarrow k$ transition, \sec^{-1}
$\sum_{\mathbf{k}} A_{\mathbf{j}\mathbf{k}}$	total transition probability for state j, sec-1
$\overline{\Lambda}_{\mathbf{j}\mathbf{k}}$	relative transition probability, $^{A}_{jk}/^{A}_{j,TOTAL}$
\overline{A}_{\searrow} .	relative transition probability at wavelength λ
D	atom orbital angular momentum = 2
d	electron orbital angular momentum = 2
$\mathbf{E}_{\mathbf{p}}$	primary electron energy (Eq. (8)), eV
Hg I,II,iII	mercury atom, single ion and double ion, respectively
J	total atom angular m
J_{B}	ion beam current, a
m	denotes Maxwell average
Nj	rate of collisionally populating the j th state from the ground
N _{jk} ,	number of j-k transitions, cm ⁻³ sec ⁻¹
N _m	Maxwell electron density, cm ⁻³
27	

neutral atom density, cm⁻³

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· N _λ	spectral line amplitude at wavelength λ
N+,++	density of singly (and doubly) ion-ized atoms, cm ⁻³
P	atom orbital angular momentum = 1
Q _j	excitation cross section, cm ²
Q _{jk}	optical excitation function for $j \rightarrow k$ transition, cm ²
r	radial position relative to thruster axis
$S_{\mathbf{j}}(\mathbf{E}_{\mathbf{p}})$	primary electron excitation coefficient at energy $\rm E_p$ to state j, cm 3 sec $^{-1}$
S _j (T _e)	Maxwell electron excitation coefficient at temperature τ_e to state j, $\mathrm{cm}^3~\mathrm{sec}^{-1}$
S	atom orbital angular momentum = 0
s	electron orbital angular momentum = 0
s_{jk}	optical excitation coefficient for $j \rightarrow k$ transition, cm ³ sec ⁻¹
$^{\rm S}\lambda$	Maxwell averaged optical excitation coefficient at wavelength λ , cm ³ sec ⁻¹
Тe	electron temperature, eV
^v e	electron speed, cm sec ⁻¹
v _p	cathode-pole piece plasma potential, v
ΔV_{I}	discharge chamber potential difference, v
$X_{jk}(E_p, T_e)$	optical excitation coefficient ratio $S_{jk}(E_p)/S_{jk}(T_e)$
$X_{\lambda}(E_{p},T_{e})$	excitation coefficient ratio at wavelength $\;\lambda\;$
φ	primary electron fraction, v_p/v_m
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primary electron density, cm⁻³

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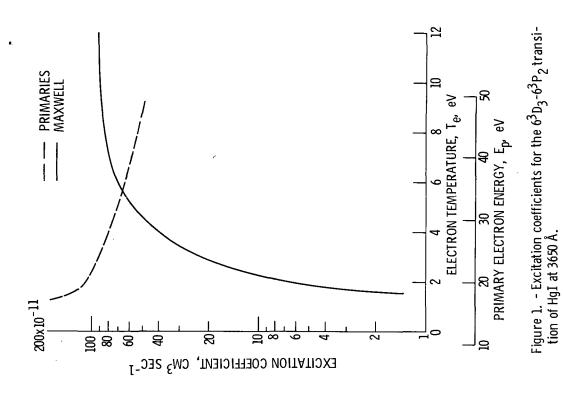
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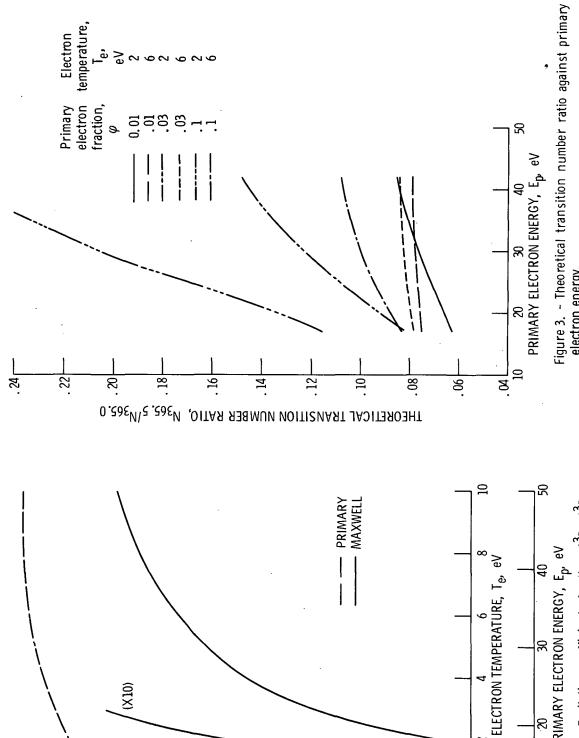
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TABLE I. - CALCULATED ION FRACTIONS

TABLE 1 CALCULATED ION FRACTIONS	Density ratios	°N/++N	<pre></pre>	0.00 - 0.007 0.006 - 0.007 0.006	0.001 .004 .004 .013	<pre></pre>
	Densit	$N_{+}/N_{o}\overline{A}(39R4)$ $(x10^{+1})$	5.5 - 8.8 7.6 - 10.5 7.8 - 8.2 7.2 - 7.1	h.5 - 7.2 $10.4 - 12.0$ 5.4 $6.9 - 8.3$	7.3 8.3 11.11 - 5.9 6.2 - 6.3	7.7 7.0 8.3 h.2 - h.h
	Primary electron	energy, eV	17 22 28 42	17 22 28 42	17 22 28 12	17 22 28 42
TAB	Radial location,	norma_1zeq	0	0.25	0.50	0.75





temperature,

(X10)

8

EXCITATION COEFFICIENT, CM^3 SEC- 1

 400×10^{-12}

200

electron energy. Figure 2. - Excitation coefficients for the $6^3 D_2 - 6^3 P_2$ transition of HgI at 3655 Å. PRIMARY ELECTRON ENERGY, Ep. eV

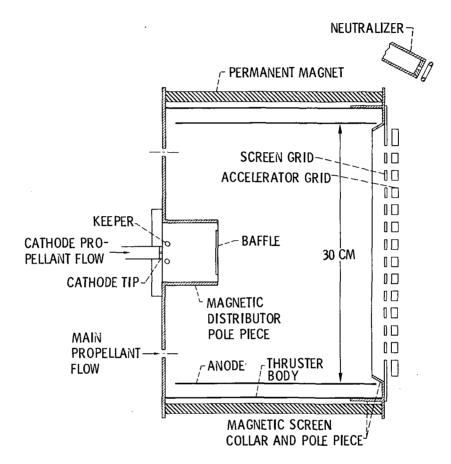


Figure 4. - Section view of a 30-centimeter-diameter thruster.

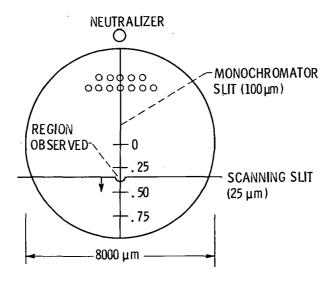


Figure 5. - 30-Centimeter-diameter thruster image viewed from downstream of the extraction grids, 72 percent open area.

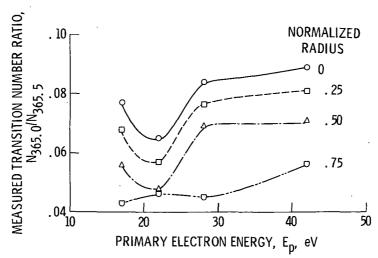


Figure 6. - Measured transition number ratio against primary electron energy, where E $_p$ = $\Delta V_{\ I}$ - 16 eV.

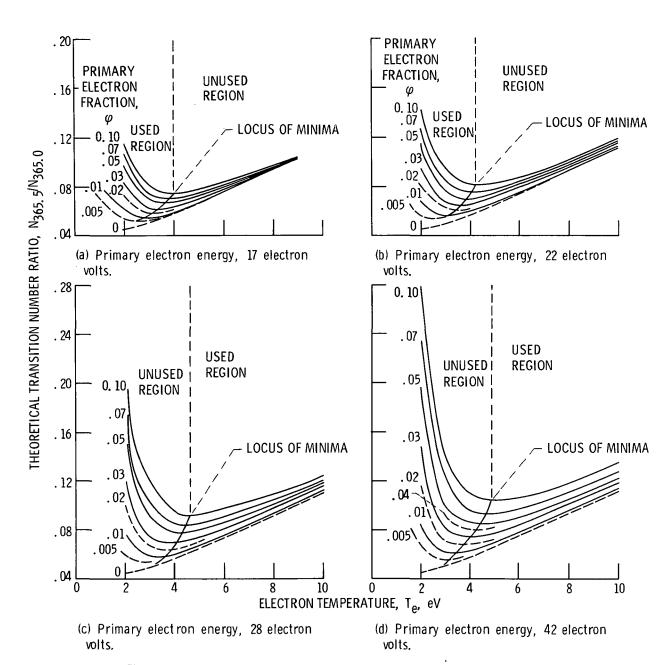
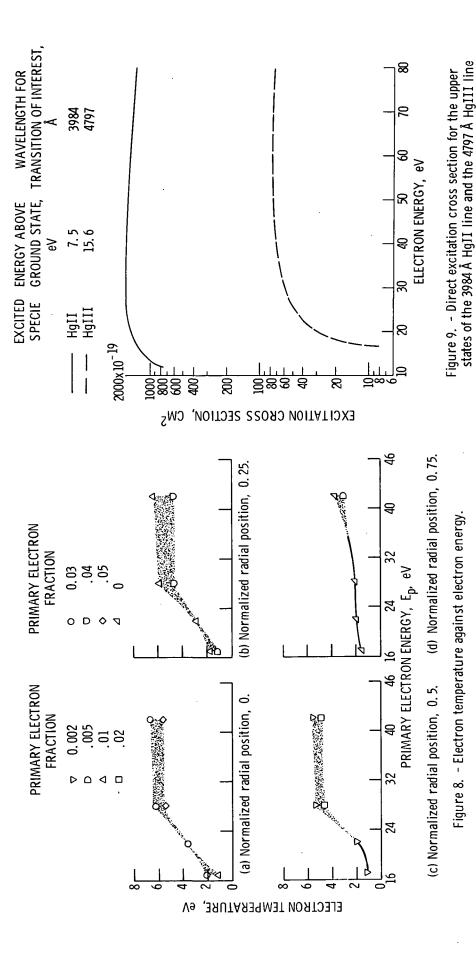


Figure 7. - Theoretical transition number ratio against electron temperature.



calculated from the Gryzinski theory. Excitation from

the ion ground states.

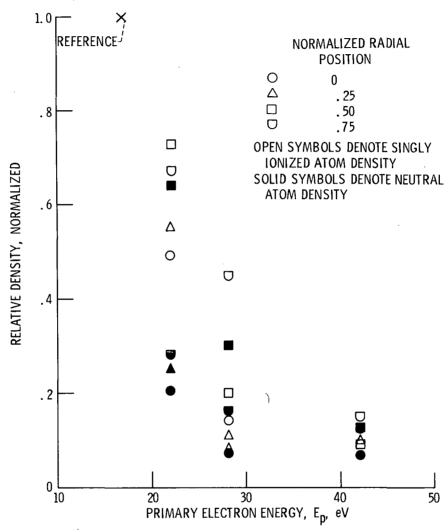


Figure 10. - Normalized ion and atom density variations with electron energy. 30-Centimeter-diameter hollow-cathode thruster; ion beam current, 1.5 amperes.